

Neutrinoless double beta decay with CUPID-Mo

Pia Loaiza

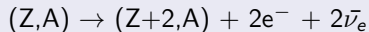


ANDES workshop, Sao Paulo, 4-6 August 2018

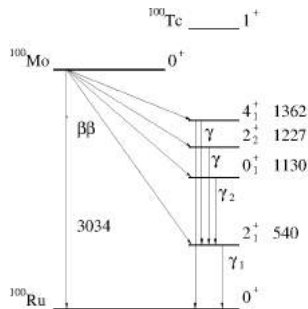
Two neutrino double beta decay

Certain isotopes are forbidden from decaying through standard beta decay because $m(Z,A) < m(Z+1,A)$

$2\beta 2\nu$

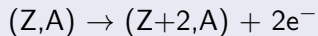


- Allowed in the **Standard Model**
- Observed for 12 isotopes
- $T_{1/2}^{2\beta 2\nu} \sim 10^{19-21}$ ys
- Important constraint for nuclear matrix elements calculation



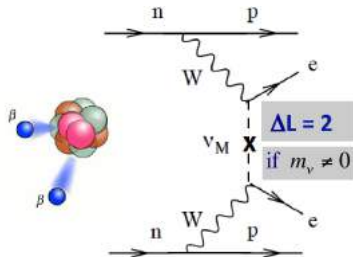
Neutrinoless Double Beta decay

$2\beta 0\nu$



If the neutrino is a **Majorana particle**, then the process of zero-neutrino double beta decay should be observable

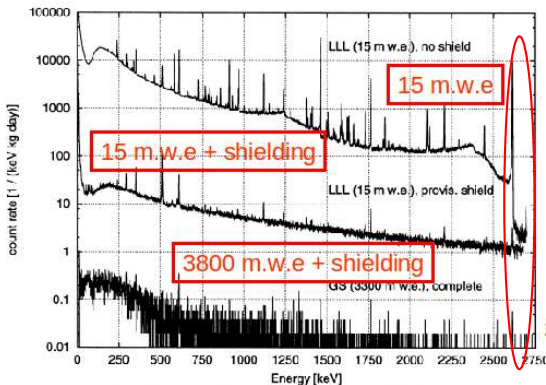
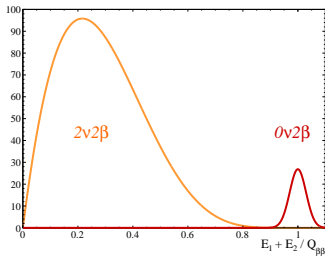
- Experimentally **not** observed
- **Implies lepton number violation**
- Offers strong support for the explanation of baryon asymmetry via leptogenesis
- Current bounds $T_{1/2}^{2\beta 0\nu} > 10^{24-26}$ ys



Experimental signature for $2\beta 0\nu$ decay

Background

Signal

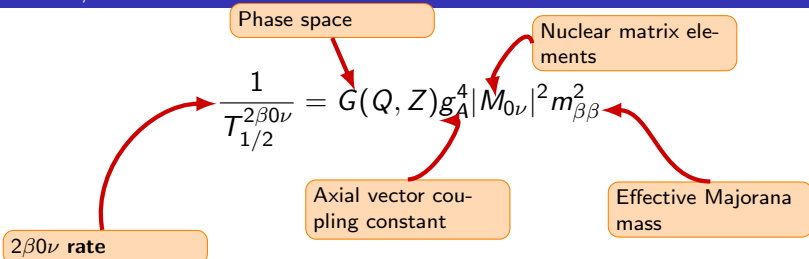


(Applied Rad and Isotopes 53 (2000) 191)

- The signal is a peak at the $Q_{\beta\beta}$ value
- The most energetic γ line from natural radioactivity is at 2615 keV \rightarrow $Q_{\beta\beta} > 2 - 3$ MeV for most promising isotopes

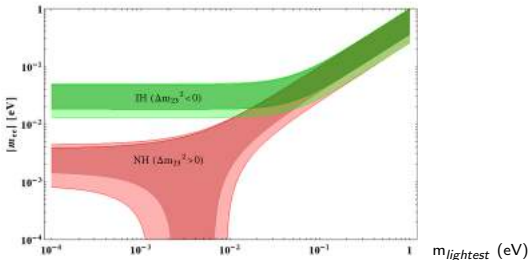
$2\beta 0\nu$ experiments measure decay rates, $T_{1/2}^{2\beta 0\nu}$

How $T_{1/2}^{2\beta 0\nu}$ is connected to neutrino masses?



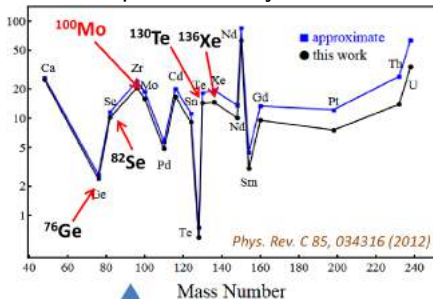
(In case of process induced by light ν exchange, mass mechanism)

$$m_{\beta\beta} = m_1|U_{e1}|^2 + m_2|U_{e2}|^2e^{i\alpha} + m_3|U_{e3}|^2e^{2i\beta}$$

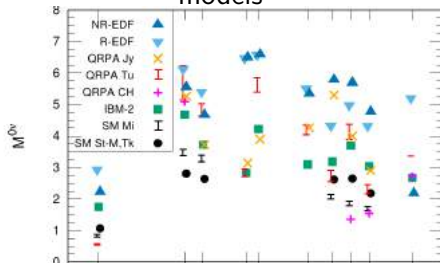


Theoretical ingredients

Phase space, exactly calculable:



Nuclear matrix elements, several models



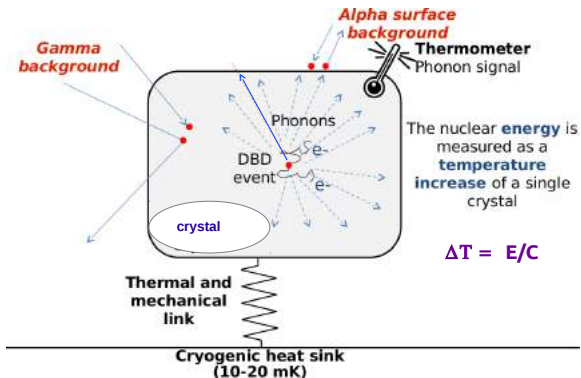
J. Engel and J. Menéndez, Rep. Prog. Phys. 80, 046301

g_A is quenched in $2\beta 2\nu$ decay. Is the renormalization the same for $2\beta 0\nu$?

- $g_A = 1.269$ free nucleon (no quenching)
- $g_A, \text{eff} \sim 0.6 - 0.8$

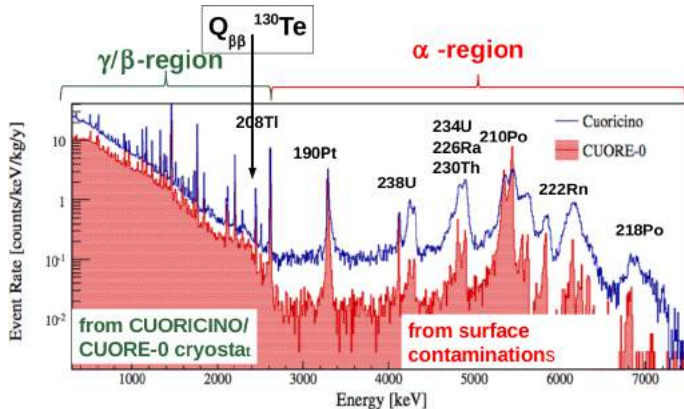
If quenching exists, the sensitivity to $m_{\beta\beta}$ will decrease

Bolometers



- **High energy resolution**
(5 keV FWHM, $\sim 0.2\%$, at $2\beta 0\nu$ ROI)
- 0.2 - 0.5 kg each crystal \rightarrow scalability to a ton scale array
- **High efficiency** ($\sim 70 - 90\%$)

Lessons learnt from CUORE

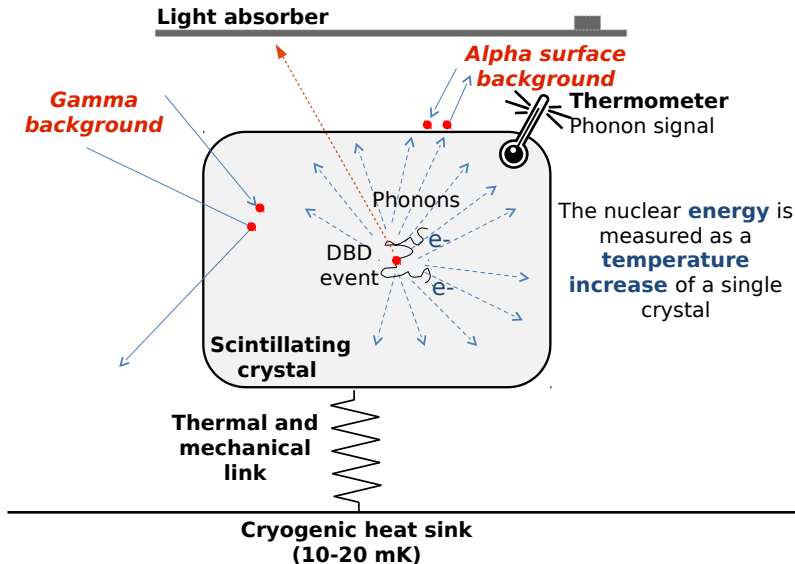


Irreducible background due to α particles emitted at the surfaces and degraded in energy

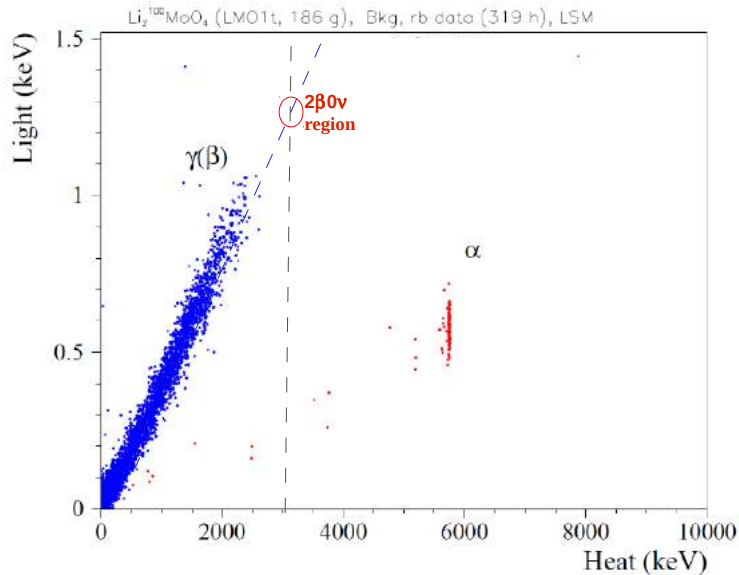
$$b \sim 10^{-2} \text{ counts/kev kg y}$$

Current solution: scintillating bolometers

Scintillating bolometers



α discrimination in a scintillating bolometer



EDELWEISS-III cryogenic facility at LSM (France)

Laboratoire Souterrain de Modane

1.7 km rock overburden (~ 4.8 km w.e.)
 $5 \mu\text{/day/m}^2$; 10^{-6} n/day/cm 2 (>1 MeV)
Deradonized air flow (~ 30 mBq/m 3)

EDELWEISS set-up

Clean room (ISO Class 4)

$^3\text{He}/^4\text{He}$ inverted wet cryostat

Passive shield

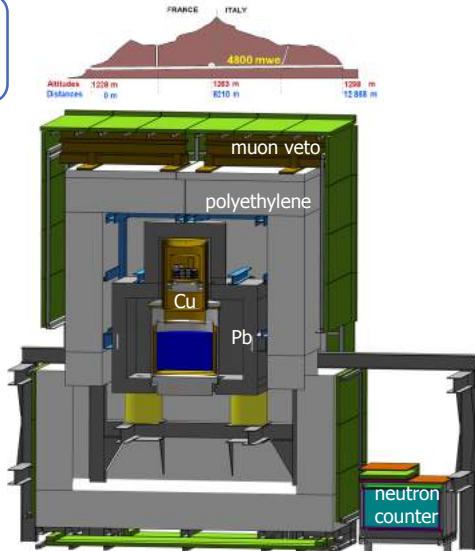
Modern lead (18 cm)
Roman lead (2 cm; 14 cm at 1 K plate)
Polyethylene (external $\sim 50+5$ cm and 10 cm at 1 K plate)

Background monitors

Muon veto (98.5% covering)
Neutron counter
Radon counter

Electronics, DAQ (Samba)

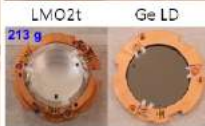
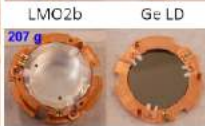
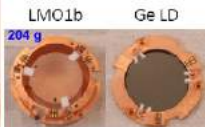
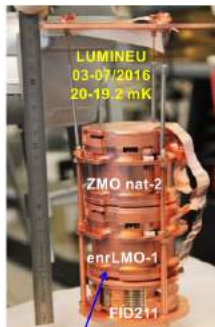
Low noise cold electronics
AC bias, modulation (100 kHz)
→ demodulation (up to 1 kHz)
16-bit or 14-bit ADC
Trigger and/or Stream data



PLB 702 (2011) 329; JINST 12 (2017) P08010; EPJC 77 (2017) 785

Tests of $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers

- Multiple tests with natural and enriched crystals in LSM and LNGS
- Longest run with 4 LMO crystals in the Edelweiss cryostat using Edelweiss electronics and DAQ (November 2016 - April 2017) [*AIP Conf. Proc. 1894, 020017 (2017)*]

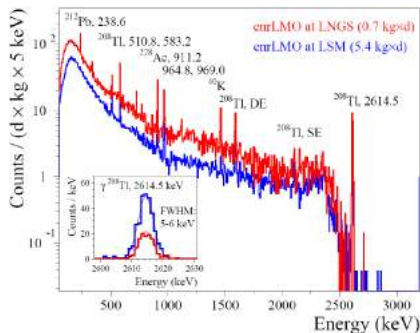
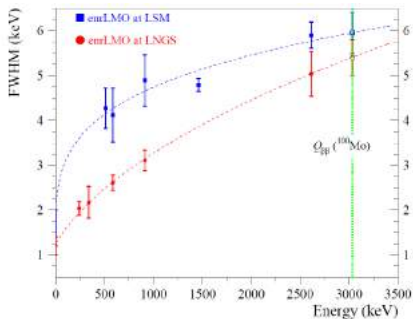


+ enrLMO-2
was tested at LNGS
05/2016, 12 mK

+ a tower T2
with enrLMO-1

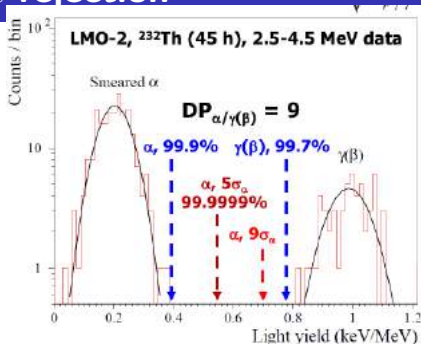
Energy resolution

Measured with a Th source (mixed ^{232}Th and ^{238}U) which allows to have several points for energy calibration



- The energy resolution (5 keV FWHM at $Q_{\beta\beta}$) required to build a next generation $2\beta0\nu$ experiment is achieved

α rejection



AIP Conf. Proc. 1894, 020017 (2017)

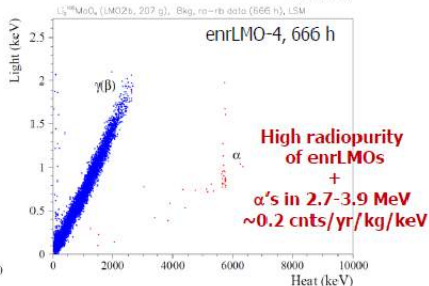
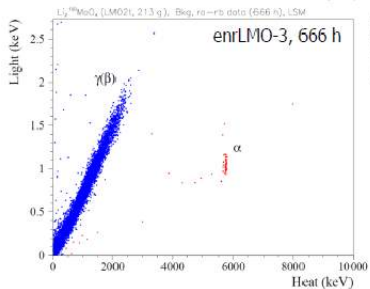
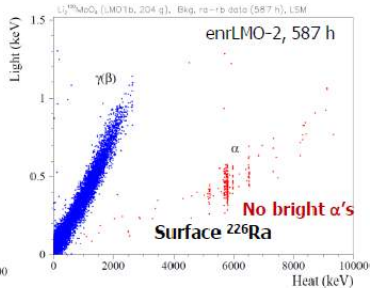
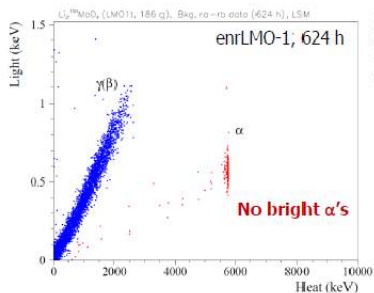
$$DP = \frac{|\mu_{\beta/\gamma} - \mu_{\alpha}|}{\sqrt{\sigma_{\beta/\gamma}^2 + \sigma_{\alpha}^2}}$$

Detector	FWHM (keV) at 2615 keV	Light Yield $_{\gamma/(\beta)}$ (keV/MeV)	$\alpha/\gamma/(\beta)$ separation above 2.5 MeV
enrLMO-1	5.8(6)	0.41	9σ
enrLMO-2	5.7(6)	0.38	9σ
enrLMO-3	5.5(5)	0.73*	$14\sigma^*$
enrLMO-4	5.7(6)	0.74*	$14\sigma^*$

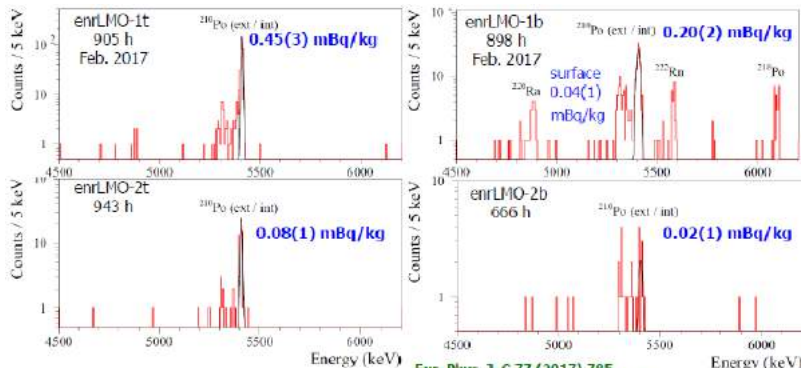
* with reflecting foil

Rejection of α 's at the level of 9σ

First measurements with 4 bolometers



Crystal radiopurity



Eur. Phys. J. C 77 (2017) 785

AIP Conf. Proc. 1894 (2017) 020017

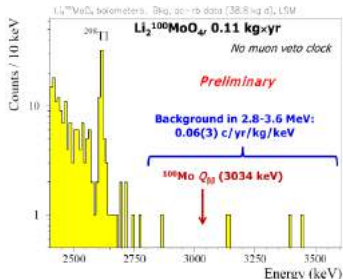
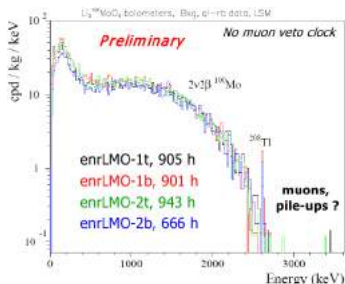
D. Poda

2

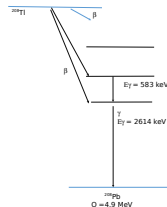
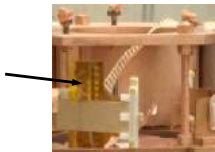
- $^{228}\text{Th} < 3 \mu\text{Bq/kg}$
- $^{226}\text{Ra} < 3 \mu\text{Bq/kg}$
- $^{210}\text{Po} : [20 - 450] \mu\text{Bq/kg}$

High radiopurity of $\text{Li}_2^{100}\text{MoO}_4$ crystals \rightarrow no background in $2\beta 0\nu$ region from internal contamination

Gamma/Beta background



Pile-ups from



Background above 2.8 MeV :
 0.06 ± 0.03 cts/(keV kg y)
 compatible with Th contamination
 from connectors close to the detector

Connectors and cabling were changed for CUPID-Mo
 Full estimation of background in progress
 Reasonable expectation : $b \sim 10^{-2} - 10^{-3}$ cts/keV kg y

CUPID-Mo demonstrators

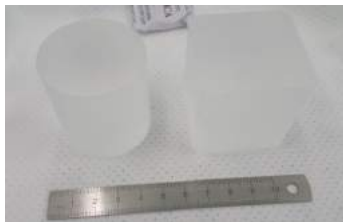
Phase I:

- 20 cylindrical $\text{Li}_2^{100}\text{MoO}_4$ crystals
→ $\sim 2.5 \text{ kg of } ^{100}\text{Mo}$
- Edelweiss set up at LSM
- **Start physics data taking end July 2018**



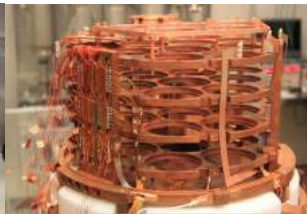
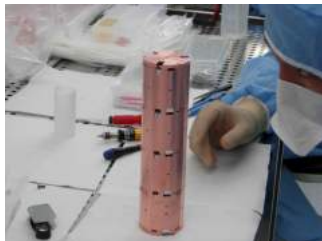
Phase II:

- Additional 26 cubic crystals
(20 + 26 cryst.) →
 $\sim 5 \text{ kg of } ^{100}\text{Mo}$
- CUPID-0 set up at LNGS
- **Planned start data taking mid-2019**

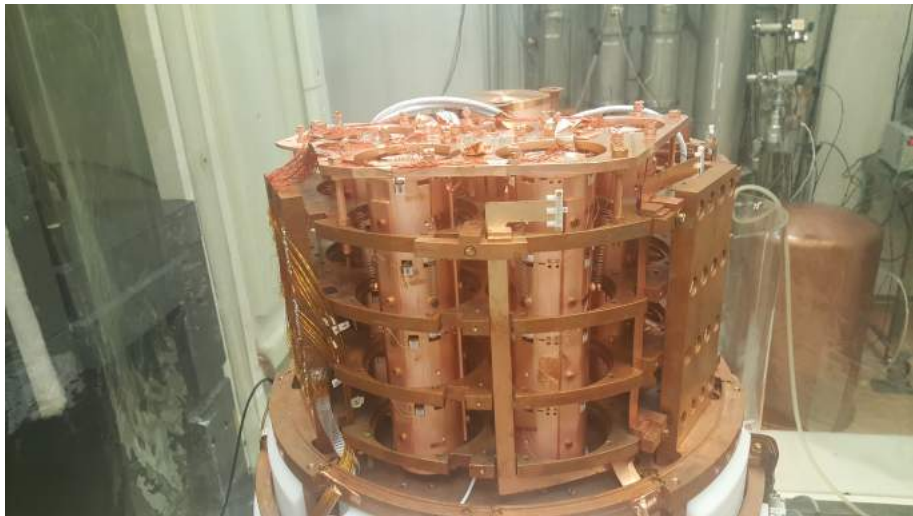


CUPID-Mo Phase I

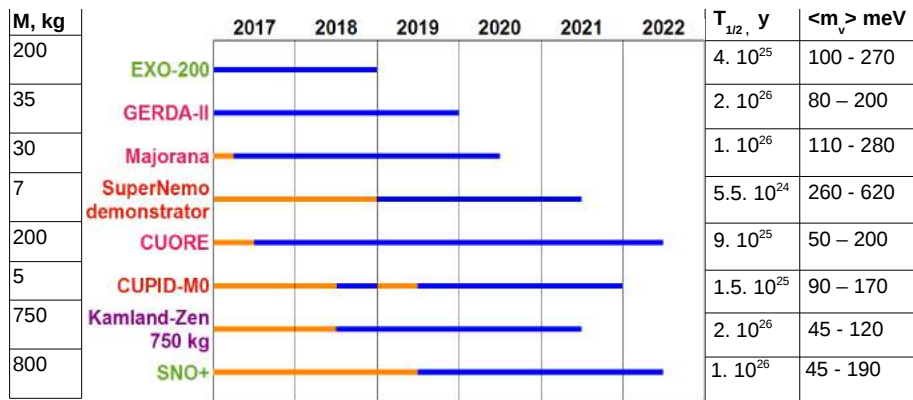
- $\text{Li}_2^{100}\text{MoO}_4$ crystals diam 44 x 45 mm
- Light detectors Ge wafer, diam 44.5 mm x 70 μm
- NTD temperature sensors
- Copper holders radiopure NOSV copper
- Spacers PTFE
- Ball bonding, 25 μm gold wires



CUPID-Mo in Edelweiss set-up



What's the near future? A possible scenario



— construction
— data

+ NEXT, AMORE, PandaX-III

CUPID-Mo : Competitive results with only 5 kg of ^{100}Mo
($b=10^{-3}$ cts/keV kg y, 10 keV ROI, 70 % eff)

Adapted from A. Barabash 'Brief review on double beta decay experiments', arXiv:1702.06340

- $g_A = 1.27$
- Phase-space factors from :

[12] Kotila J and Iachello F 2012 *Phys. Rev. C* **85** 034316.

[13] Mirea M Pahomi T and Stoica S 2015 *Rom. Rep. Phys.* **67** 879.

- NME from :

[3] Hyvarinen J and Suhonen J 2015 *Phys. Rev. C* **91** 024613.

[4] Simkovic F Rodin V Faessler A and Vogel P 2013 *Phys. Rev. C* **87** 045501.

[5] Barua J Kotila J and Iachello F 2015 *Phys. Rev. C* **91** 034304.

[6] Rath P K *et al.* 2013 *Phys. Rev. C* **88** 064322.

[7] Rodriguez T R and Martinez-Pinedo G 2010 *Phys. Rev. Lett.* **105** 252503.

[8] Menendez J Poves A Caurier E and Nowacki F 2009 *Nucl. Phys. A* **818** 139.

[9] Neacsu A and Horoi M *et al.* 2015 *Phys. Rev. C* **91** 024309.

[10] Mustonen M and Engel J 2013 *Phys. Rev. C* **87** 064302.

[11] Song L S *et al.* 2017 *Phys. Rev. C* **95** 024305.

CUPID: Cuore Upgrade with Particle IDentification

Follow-up of CUORE, towards a **ton scale bolometric experiment** with a factor 100 background reduction.

R&D efforts in three axis:

- 1 $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers
- 2 $^{130}\text{TeO}_2$ Cherenkov bolometers
- 3 Zn^{82}Se scintillation bolometers

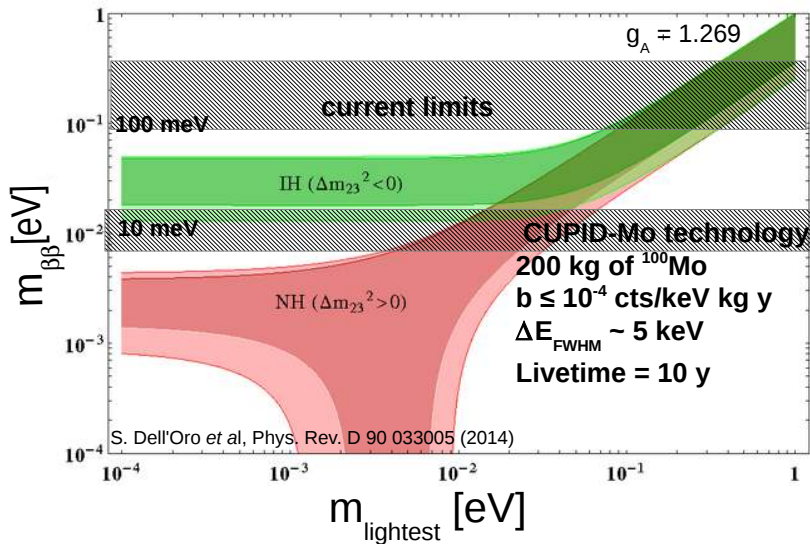
Based on the result of on-going R&D and demonstrator experiments, **$\text{Li}_2^{100}\text{MoO}_4$ is identified as a promising baseline** and $^{130}\text{TeO}_2$ Cherenkov as a mature viable alternative.

Purpose: fully explore the $m_{\beta\beta}$ Inverse Hierarchy region

- $\mathbf{b} \sim 10^{-4}$ counts/kev kg y
- $T_{1/2} > 10^{27}$ y

CUPID collaboration will be formed in the near future
CUPID kick-off meeting planned fall 2018

Next generation experiments



Summary

- Study of neutrinoless double beta decay is one of the most urgent topics in particle physics and cosmology
- The bolometric approach is a viable technique confirmed at large scale by the CUORE results
- A promising technology based on enriched $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers was developed and is now applied to the CUPID-Mo demonstrator
- CUPID (Cuore Upgrade with Particle IDentification) is one of the most promising next-generation searches

CUPID-Mo collaboration

Follow up of LUMINEU collaboration (ANR-French funding, 2012-2017)

<http://cupid-mo.mit.edu>



7 countries, 15 institutions, ~110 scientists



- CSNSM Orsay, CEA/DRF Gif-sur-Yvette, IPNL Lyon, LAL Orsay, FRANCE
- KIT Karlsruhe, GERMANY
- INFN Bicocca and Roma, LNGS INFN L'Aquila, ITALY
- KINR Kyiv, UKRAINE
- JINR Dubna, ITEP Moscow, NIC Novosibirsk, RUSSIA
- MIT Boston, UCB/LBNL Berkeley, US
- CUPID-China: Fudan Shanghai, USTC Hefei, CHINA

EXTRA SLIDES

Prospect for next generation experiments

Experiment	Mass (kg)	t (y)	Sensitivity $T_{1/2}$ (y)	Sensitivity $\langle m_\nu \rangle$ meV
CUPID	200	10 (2022? - 2032)	$2.2 \cdot 10^{27}$	6 - 17
nEXO	5000	10 (2025? - 2035)	$10^{27} \cdot 10^{28}$	6 - 53
LEGEND	200	4 (2022? - 2026)	$1.0 \cdot 10^{27}$	34 - 91
KamLAND-Zen	1000	3 (2020? - 2023)	$6.0 \cdot 10^{26}$	26 - 69
SNO+	8000	5 (2020?- 2025)	$7 \cdot 10^{26}$	20-73
SuperNEMO	100	?	$1.0 \cdot 10^{26}$	40-140

For CUPID-Mo:

Assuming $b=10^{-4}$ counts /keV kg y, 10 year running, 8 keV energy window, 78% efficiency